



UNIVERSITY OF ICELAND

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ENTANGLEMENT ENTROPY  
AT NON - EQUILIBRIUM  
IN HOLOGRAPHY

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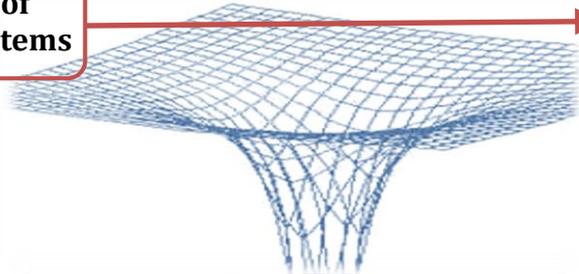
Based on ArXiv: **1705.04696**, in collaboration with

- Johanna Erdmenger (University of Würzburg)
- Mario Flory (Jagiellonian University of Kraków)
- Eugenio Megías (University of the Basque Country)
- Ann-Kathrin Straub (Max Planck Institute for Physics, Munich)
- Piotr Witkowski (Max Planck Institute for Complex Physical Systems, Dresden)



# CONTEXT

Real time dynamics of strongly correlated systems



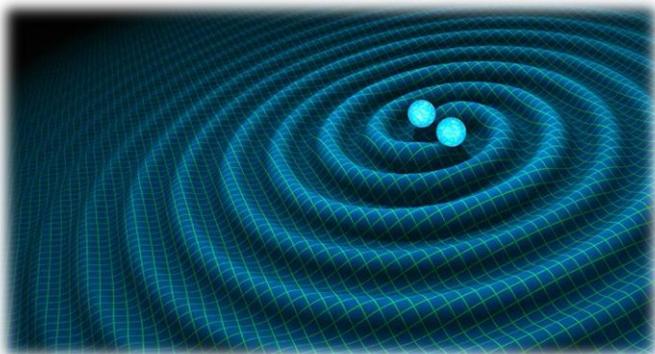
- Directly computable
- Easy collective responses

- ❖ Quark-gluon plasma thermalization [*Chesler, Yaffe, Heller, Romatschke, Mateos, van der Schee*]
- ❖ Quantum quenches [*Balasubramanian, Buchel, Myers, van Niekerk, Das*]
- ❖ Driven superconductors [*Rangamani, Rozali, Wong*]



*Important conclusion:*

Transition to hydrodynamic regime occurs very early!



- ❖ Turbulence in Gravity [*Lehner, Green, Yang, Zimmerman, Chesler, Adams, Liu*]



Insight into gravity gained from high-energy physics

# MOTIVATION

**Emergent Collective Behavior:** Quantum effects ↔ Out of equilibrium physics



Context: Quantum Mechanics of many-body systems

→ *How can we make predictions?*

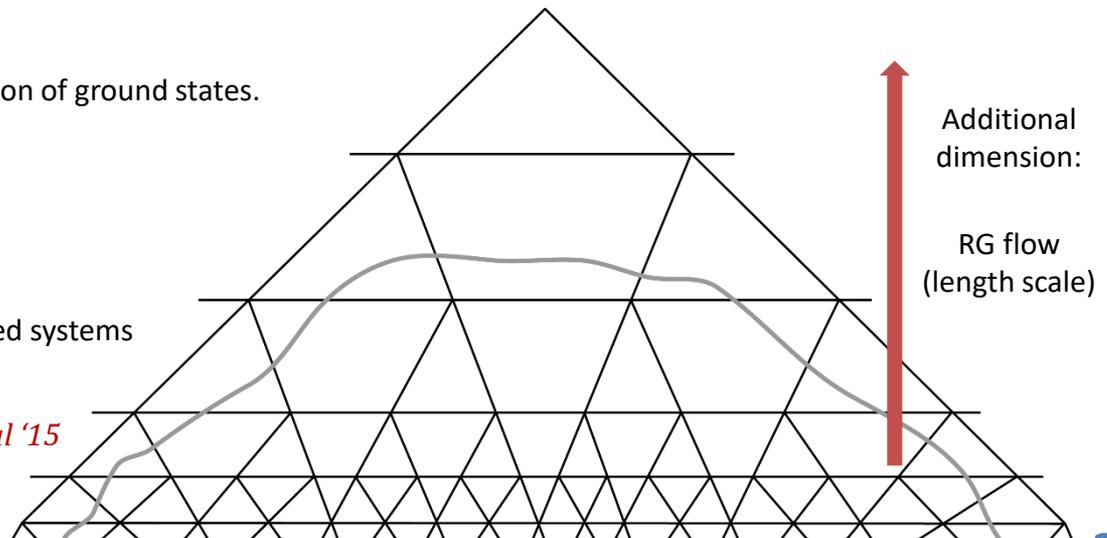
- 1) **Entanglement:** Indicates structure of global wave function.
- 2) **RG group:** Increasing length scale, a sequence of effective descriptions is obtained.
- 3) **Entanglement Renormalization:** Careful removal of short-range entanglement.

4) **Tensor Networks:** Effective description of ground states.

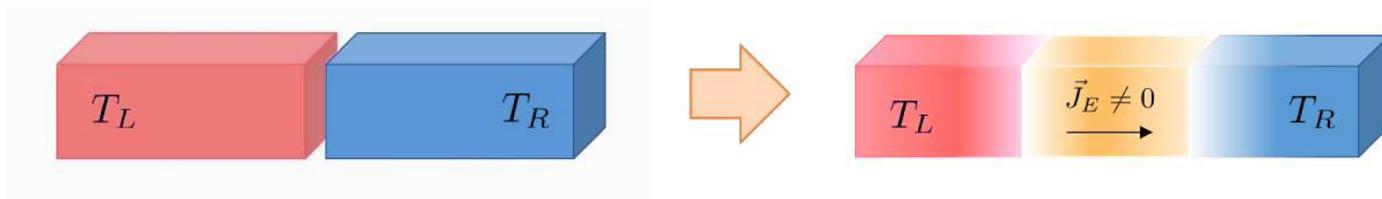


Analysis of entanglement  
to ascertain spatial structure of strongly coupled systems

*Evenbly, Vidal '15*



# SETUP EXPLANATION



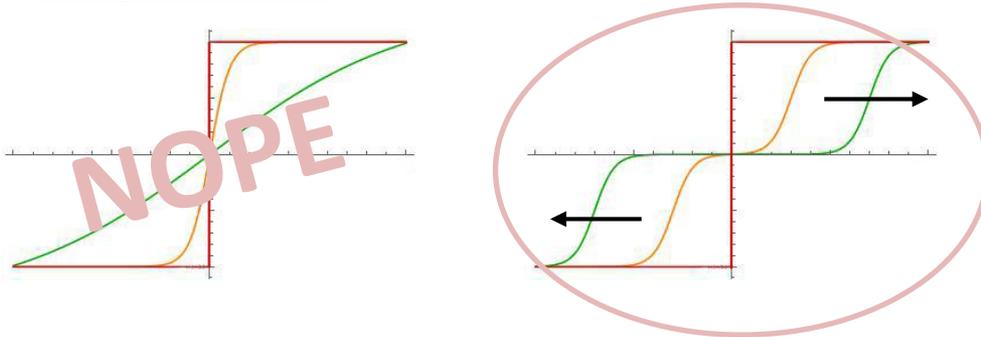
*Bernard, Doyon '12*  
*Chang, Karch, Yarom '13*

➤ Initial configuration:

1+1 dimensional system separated into **two** regions,  
**independently** prepared in thermal equilibrium.

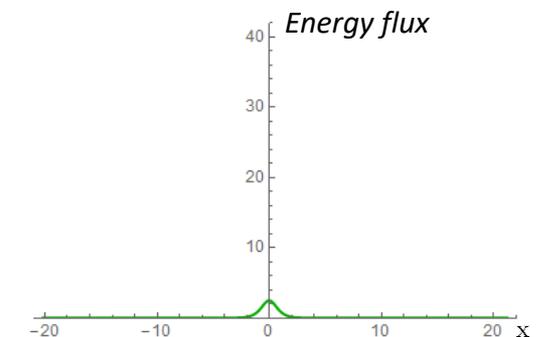
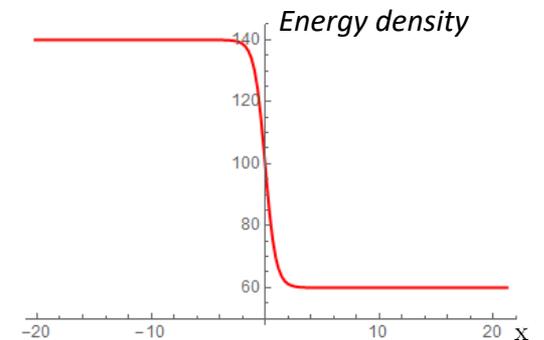
$$T(t = 0, x) = T_L \theta(-x) + T_R \theta(x)$$

➤ Subsequent evolution:



A growing region with a constant energy flow, the **steady state**, develops.  
 This region is described by a thermal distribution at shifted temperature.

The state carries a constant energy current.



# HISTORY REVIEW

*Bhaseen, Doyon, Lucas, Schalm '13*

*Bernard, Doyon '12*

## Thermal quench in 1+1

Two exact copies initially at equilibrium, independently thermalized.



Conservation equations & tracelessness:

$$\begin{aligned} \partial_x \langle T^{xx} \rangle &= -\partial_t \langle T^{tx} \rangle = 0 \\ \langle T^{xx} \rangle &= \langle T^{tt} \rangle \end{aligned}$$



$$\begin{cases} \langle T^{tx} \rangle = F(x-t) - F(x+t) \\ \langle T^{tt} \rangle = F(x-t) + F(x+t) \end{cases}$$

## **Expectation for CFT:**

Shock waves emanating from interface, converge to non-equilibrium *Steady State*.

## Generalization to any d

- Assume ctant. homogeneous heat flow as well:

$$\langle T^{\mu\nu} \rangle = a_d T^{d+1} (\eta^{\mu\nu} + (d+1)u^\mu u^\nu)$$

- Effective dimension reduction to 1+1.
- Linear response regime:

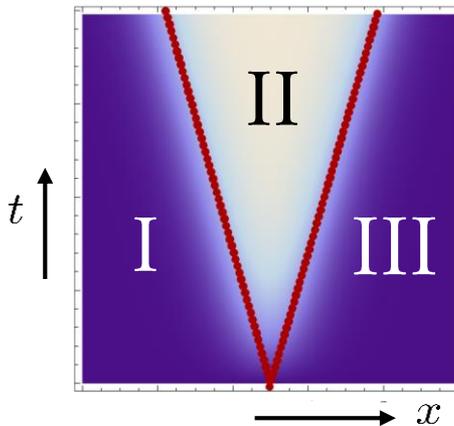
$$|T_L - T_R| \ll T_L + T_R$$

→ Hydro eqs. explicitly solvable.

*Bhaseen, Doyon, Lucas, Schalm '13*  
*Chang, Karch, Yarom '13*

## Hydrodynamical evolution of 3 regions

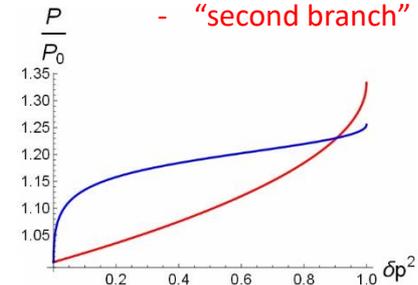
Match solutions ↔ Asymptotics of the central region.



$$T^{\mu\nu} = P(d u^\mu u^\nu + \eta^{\mu\nu}) + \pi^{\mu\nu} + \mathcal{O}(\partial^2)$$

Two configurations:

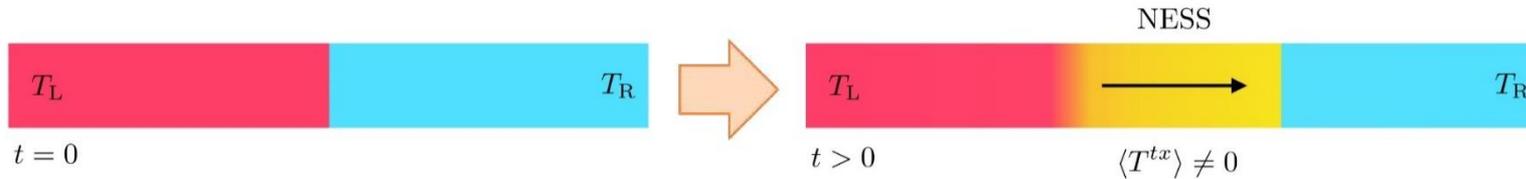
- Thermodynamic branch
- "second branch"



# RAREFACTION WAVE

- There is no uniqueness of solution to the non-linear PDEs.
  - Doble shock solution: *Mathematically correct, but not physical.*
  - New solution: shock + rarefaction.

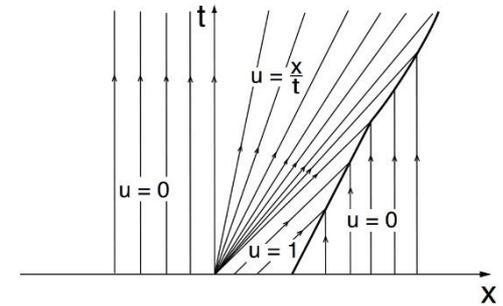
Spillane, Herzog '15  
Lucas, Schalm, Doyon, Bhasen '15  
Hartnoll, Lucas, Sachdev '16



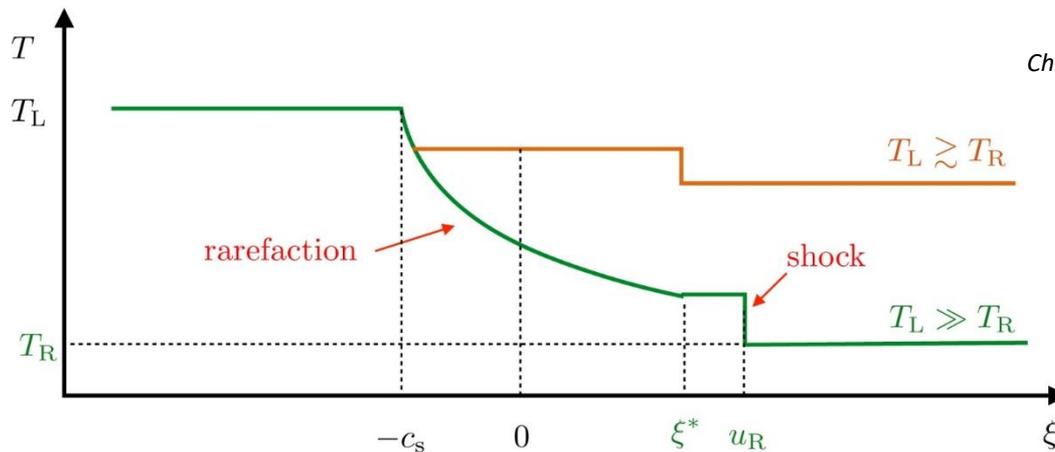
## Entropy condition

*Riemann problem:* When we have conservation equations like  $\partial_t u + \partial_x(f(u)) = 0$ , the curves along which the initial condition is transported must end on the shock wave.

- The speed of the solution must be  $f'(u_L) > (u_L + u_R)/2 > f'(u_R)$ , which rules out a shock moving into the hotter region.



Characteristics must end in the shockwave, not begin.



# ENTANGLEMENT TSUNAMIS

Liu, Suh '13

Li, Wu, Wang, Yang '13

**Context:** A global quench leading to an AdS black hole as final state.

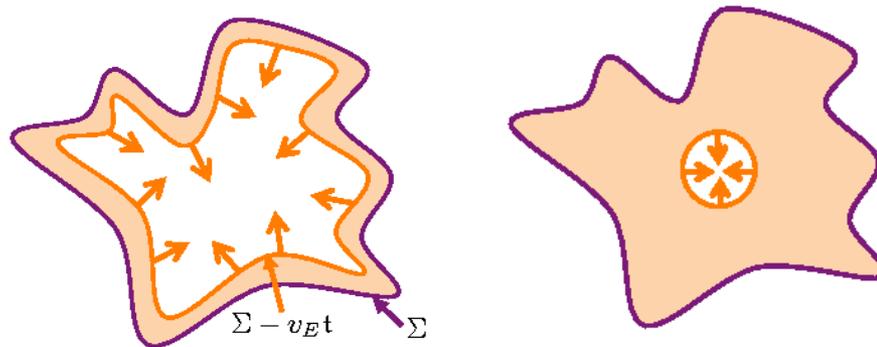
(Thin shell of matter which collapses to form a black hole)

$$ds^2 = \frac{L^2}{z^2} (-[1 - \theta(t)g(z)] dt^2 - 2dt dz + d\vec{x}^2)$$

**Entanglement growth:** Initially quadratic, then followed by a universal linear regime.

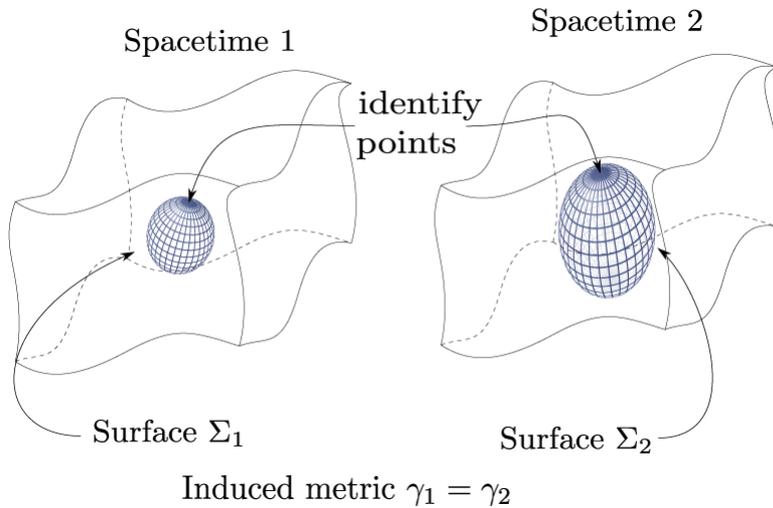
$$\Delta S_\Sigma(t) = \frac{\pi}{d-1} \mathcal{E} A_\Sigma t^2 + \dots \quad \longrightarrow \quad \Delta S_\Sigma(t) = s_{\text{eq}} (V_\Sigma - V_{\Sigma-v_E t}) t + \dots$$

*Simple geometric picture:* A wave with a sharp wave-front propagating inward from  $\Sigma$ , and the region that has been covered by the wave is entangled with the region outside  $\Sigma$ , while the region yet to be covered is not so entangled.



# GLUING SPACETIMES

Israel '66



Take two spacetimes and define codimension one hypersurfaces  $\Sigma_{1/2}$  such that they have the same topology.

If the induced metric on  $\Sigma_{1/2}$  is the same ( $\gamma_1 = \gamma_2 = \gamma$ ), the two spacetimes can be matched by identifying  $\Sigma_{1/2}$  if the energy-momentum on  $\Sigma$  satisfies

$$(K_{ij}^+ - \gamma_{ij}K^+) - (K_{ij}^- - \gamma_{ij}K^-) = -\kappa S_{ij}$$

- Israel Junction Conditions -

*In our setup:*

$$ds^2 = \begin{cases} ds_{TL}^2 & \text{if } x < -t \\ ds_{\text{boost}}^2 & \text{if } -t < x < t \\ ds_{TR}^2 & \text{if } x > t \end{cases}$$

Discontinuous geometry!

$S_{ij}$  : Energy momentum tensor on the surface  
 $\gamma_{ij}$  : Induced metric  
 $K^{\pm}$  : Extrinsic curvatures depending on embedding.

$$ds^2 = \frac{L^2}{z^2} (dz^2 + g_{\mu\nu}(z, x, t)dx^\mu dx^\nu)$$

There is an analytic solution...

$$g_{tt}(z, x, t) = - \left[ 1 - \frac{z^2}{L^2} (f_R(x-t) + f_L(x+t)) \right]^2 + \left[ \frac{z^2}{L^2} (f_R(x-t) - f_L(x+t)) \right]^2$$

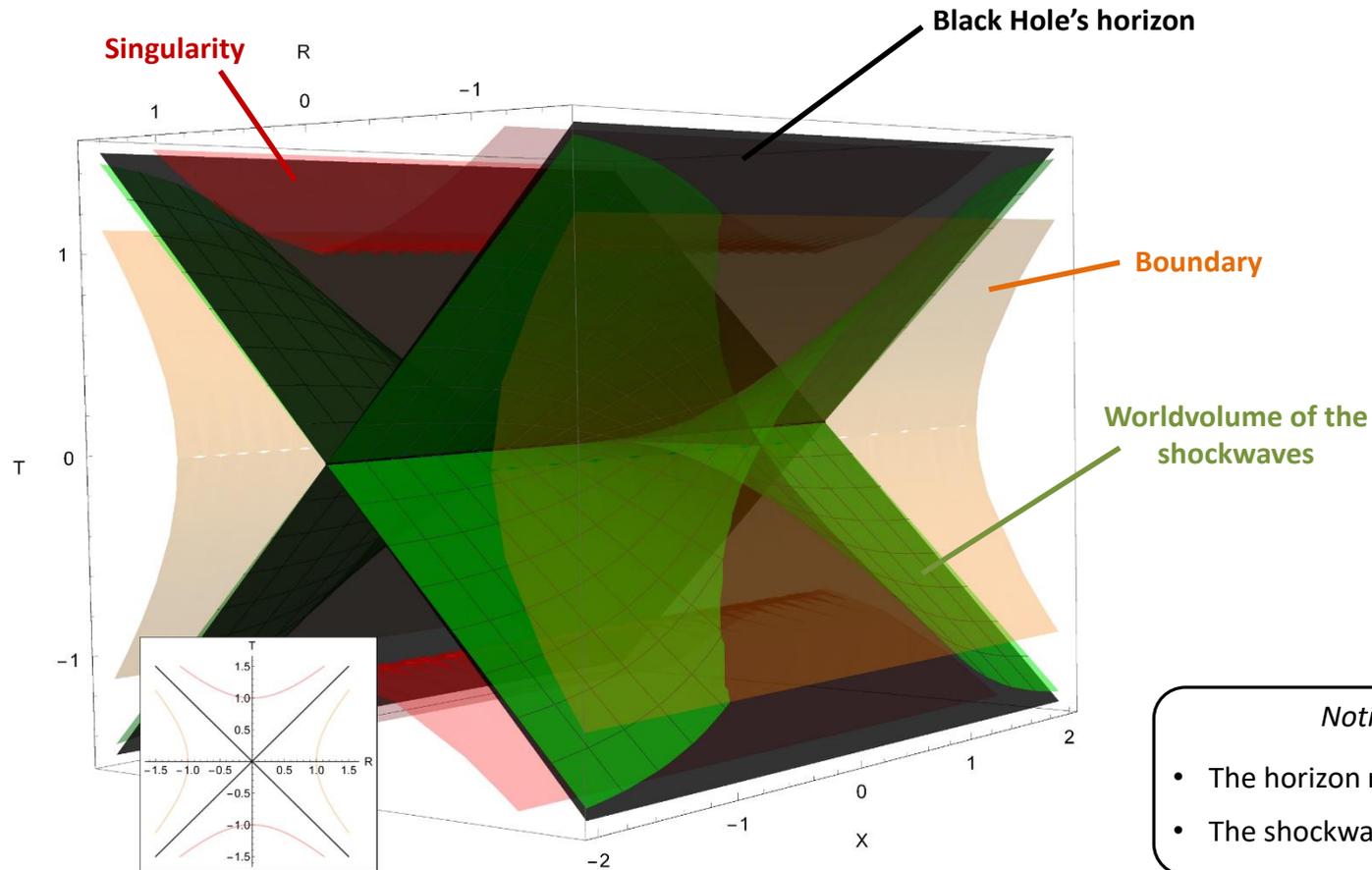
with  
 Initial condition:  $f_{L/R}(v) \rightarrow \frac{\pi^2 L^2}{2} (T_L^2 + (T_R^2 - T_L^2) \theta(v))$

But... is the horizon cut into 3 pieces??

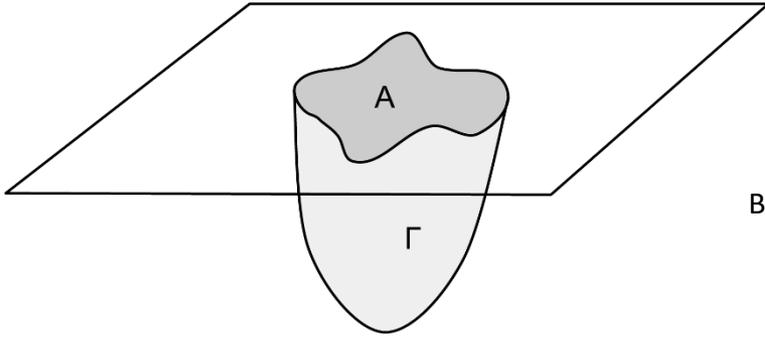
# SPACETIME DIAGRAM

Coordinates compactified:

$$T = \left| \frac{z - z_H}{z + z_H} \right|^{1/2} \sinh \left( \frac{t}{z_H} \right), \quad R = \left| \frac{z - z_H}{z + z_H} \right|^{1/2} \cosh \left( \frac{t}{z_H} \right)$$



# ENTANGLEMENT ENTROPY



The geometry is discontinuous:

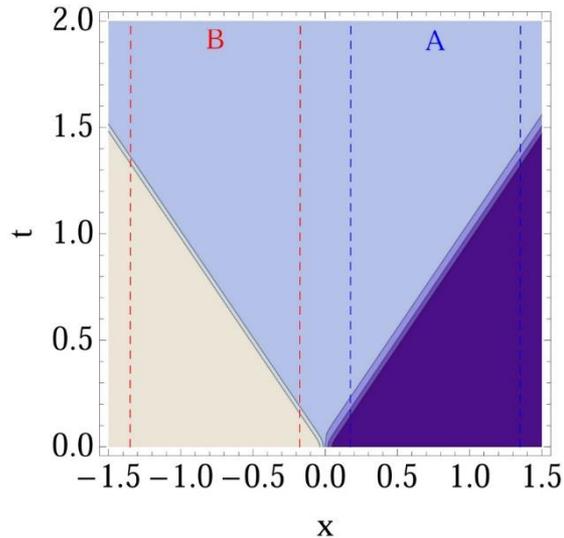
$$f_{L/R}(v) \rightarrow \frac{\pi^2 L^2}{2} (T_L^2 + (T_R^2 - T_L^2) \theta(v))$$

But we replace the step function by this initial condition

$$f_L(v) = f_R(v) = \frac{\pi^2 L^2}{4} \left( (T_L^2 + T_R^2) + (T_R^2 - T_L^2) \tanh(\alpha v) \right)$$

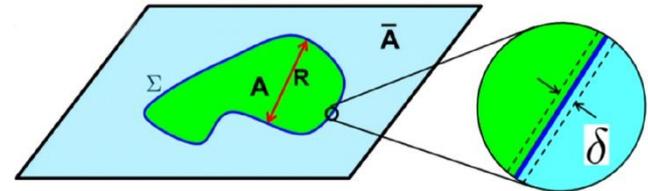
➤ **Shooting method to find geodesic lengths:** Shoot from the tip until the desired boundary values are obtained.

Intervals A, B considered:



Contour plot of the energy density

**Entanglement Regularization:**

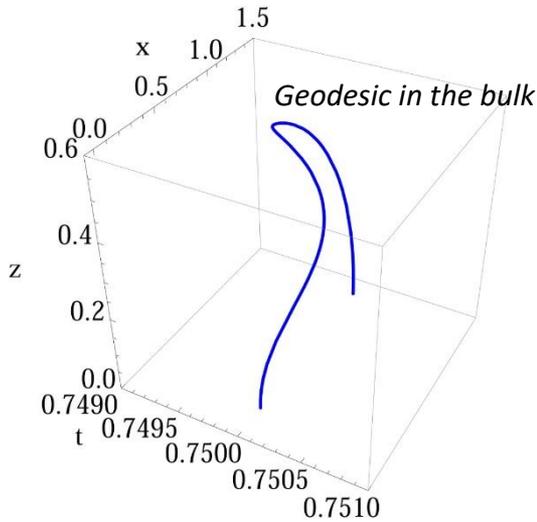


- Small distance contributions must be subtracted
- We use minimal subtraction scheme:

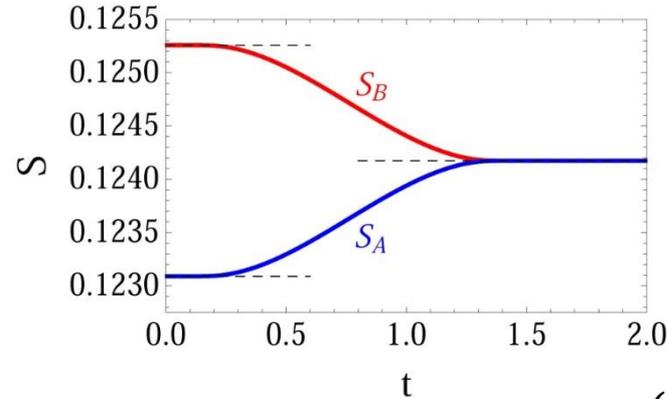
$$S_A^{\text{ren}} = \frac{1}{4G} (\text{Area}(\gamma_A) - \text{Area}(\gamma_A^{\text{div}}))$$

with  $\text{Area}(\gamma_A^{\text{div}}) = -2L \log \delta$

# UNIVERSAL LAW



Time evolution of the entanglement entropy of intervals A and B:



Define the normalized entanglement entropy:

$$f_A(\rho) \equiv \frac{S_A(t) - S_A(t=0)}{S_A(t=\infty) - S_A(t=0)}$$

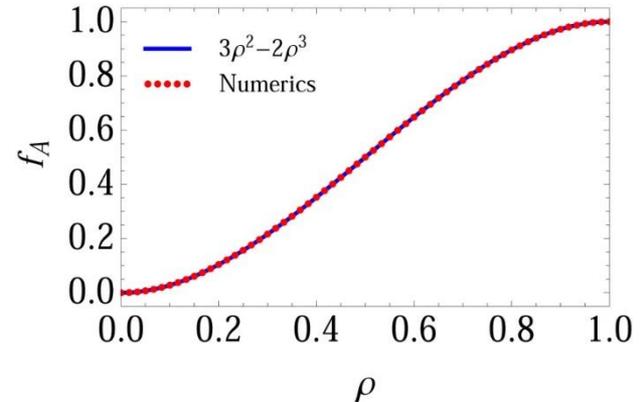
where

$$\rho \equiv (t - t_1)/\ell$$

$$S_A(t) = \begin{cases} S_A(t=0) & 0 \leq t \leq t_1 \\ S_A(t) & t_1 \leq t \leq t_2 \\ S_A(t=\infty) & t_2 \leq t \end{cases}$$

Plots overlay on top of each other,  
numerical behavior seems to be well approximated by

$$f_A(\rho) \simeq 3\rho^2 - 2\rho^3, \quad 0 \leq \rho \leq 1$$



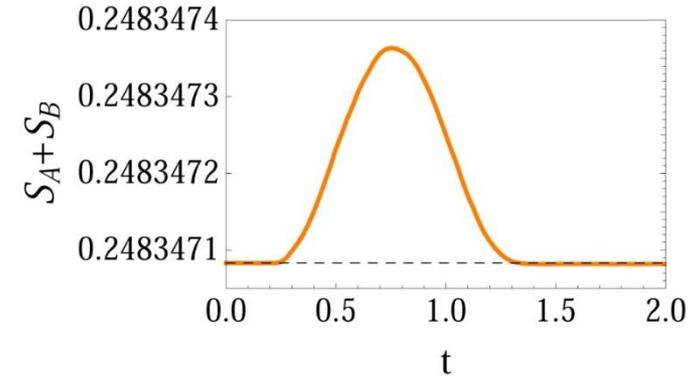
# NON-UNIVERSAL EFFECTS

❖ Conservation of entropy:

$$S_A(t=0) + S_B(t=0) = S_A(t=\infty) + S_B(t=\infty)$$

But it's not conserved at intermediate times!

$$S_{A+B}(t) \neq \text{const} \implies$$

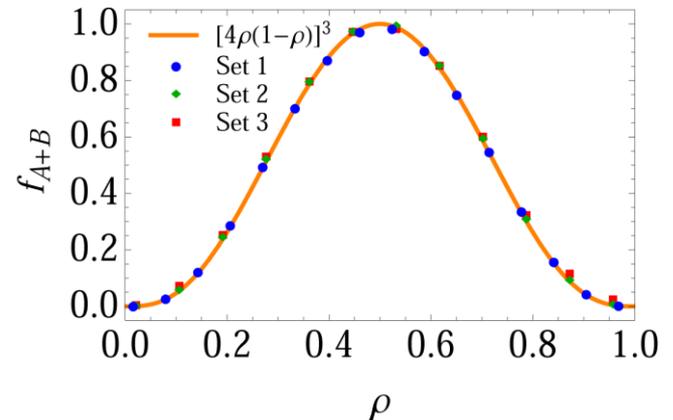


❖ Define the normalized total entanglement entropy:

$$f_{A+B}(\rho) \equiv \frac{S_{A+B}(t) - S_{A+B}(t=0)}{S_{A+B}(t_{\max}) - S_{A+B}(t=0)} \quad \text{where } \rho \equiv (t - t_1)/\ell$$

Plots overlay on top of each other,  
numerical behavior seems to be well approximated by

$$f_{A+B}(\rho) \simeq [4\rho(1-\rho)]^3, \quad 0 \leq \rho \leq 1$$



❖ **Conclusion:** Non-conservation effects are caused by non-universal contribution:

$$f_A(\rho) = 3\rho^2 - 2\rho^3 + \boxed{C(T_L, T_R, \ell)} [4\rho(1-\rho)]^3$$

Factor with non-universal dependence  
on the parameters of the interval

# MUTUAL INFORMATION

How does information get exchanged between the systems which are isolated at  $t=0$ ?

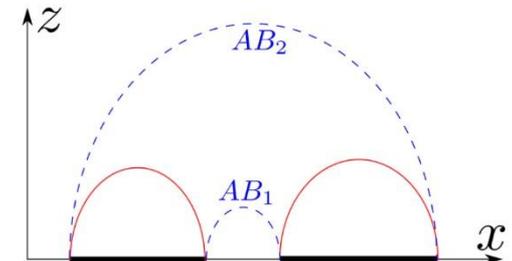
**Def.)**  $I(A, B) = S_A + S_B - S(A \cup B)$  where  $S(A \cup B) = \min \left\{ S_A + S_B, S_1 + S_2 \right\}$

Interpretation:

It measures which information of subsystem A is contained in subsystem B.

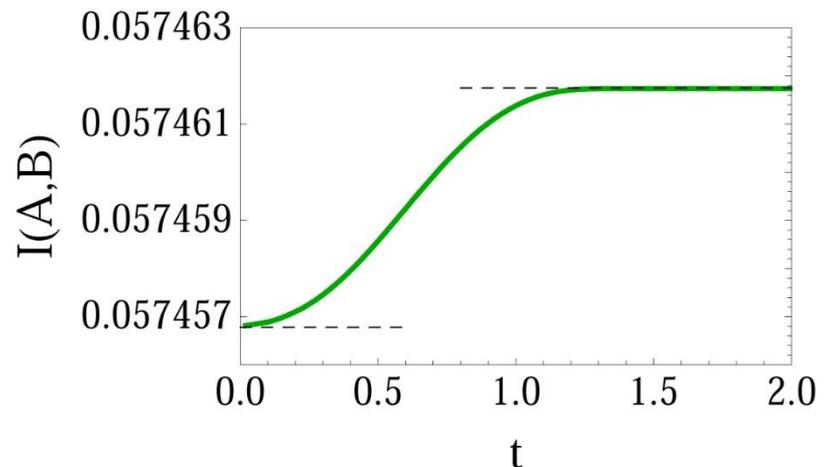
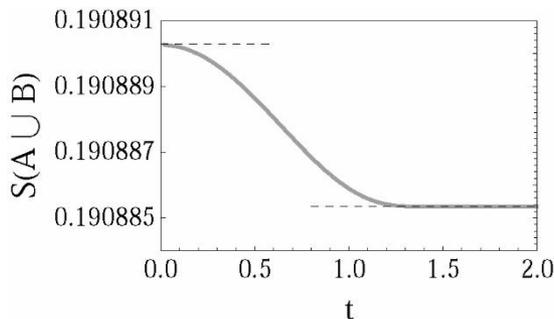
In other words: The amount of information that can be obtained from one of the subsystems by looking at the other one.

Note that  $I(A, B) \geq 0$  always.

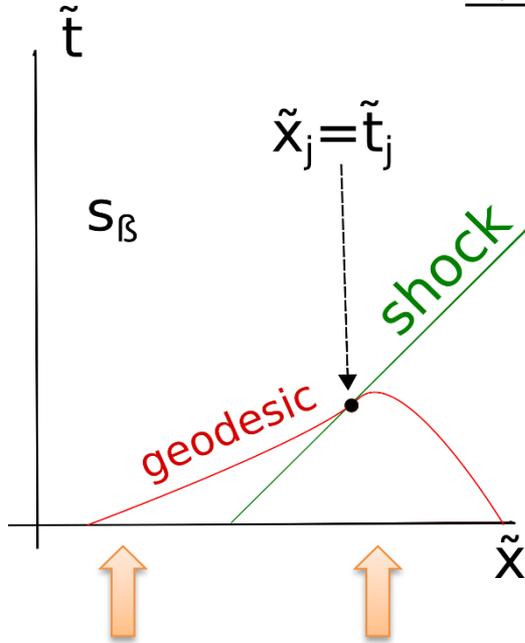


Observation:  $\partial_t I(A, B) \geq 0$

- ❖ The shockwaves transport information about the presence of the other heat bath., although they are spacelike in the bulk.



# MATCHING GEODESICS



One end on the steady state,  
another in the thermal region.

Condition for the position of the shockwave:

$$x_j = t_j \Leftrightarrow \tilde{x}_j = \tilde{t}_j$$

## Complementary approach – Steps:

- 1) Calculate geodesics in each spacetime region.
- 2) Add their renormalized lengths
- 3) Extremize the sum with respect to the meeting point.

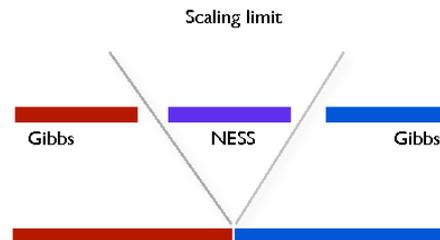
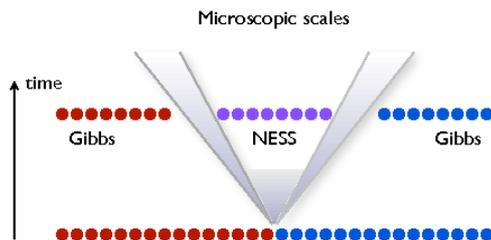
In previous calculations,

$$f_L(v) = f_R(v) = \frac{\pi^2 L^2}{4} \left( (T_L^2 + T_R^2) + (T_R^2 - T_L^2) \tanh(\alpha v) \right)$$

Here, the metric components are discontinuous

$$f_{L/R}(v) \rightarrow \frac{\pi^2 L^2}{2} (T_L^2 + (T_R^2 - T_L^2) \theta(v))$$

→ Agreement between numerical results at large  $\alpha$   
and results from this approach?



**Note:**

$$ds^2 = \frac{L^2}{z^2} (dz^2 + g_{\mu\nu}(z, x, t) dx^\mu dx^\nu)$$

→ Schwarzschild coordinates!

# SMALL TEMPERATURE EXPANSION

❖ In this limit, we can prove the previous universal law:

$$f_A(\rho) \simeq 3\rho^2 - 2\rho^3, \quad 0 \leq \rho \leq 1$$

❖ The replacement is:

$$T_L \rightarrow \delta T_L, \quad T_R \rightarrow \delta T_R \quad \longrightarrow \quad \begin{aligned} \partial_x d_R &\propto \partial_{z_j} d_L \propto (\ell - t)(\ell + 2t - 4x)t - (\ell - 2t)z_j^2, \\ \partial_{z_j} d_R &\propto \partial_x d_L \propto (\ell - t)(t - 2x)(\ell + t - 2x)t + z_j^4. \end{aligned}$$

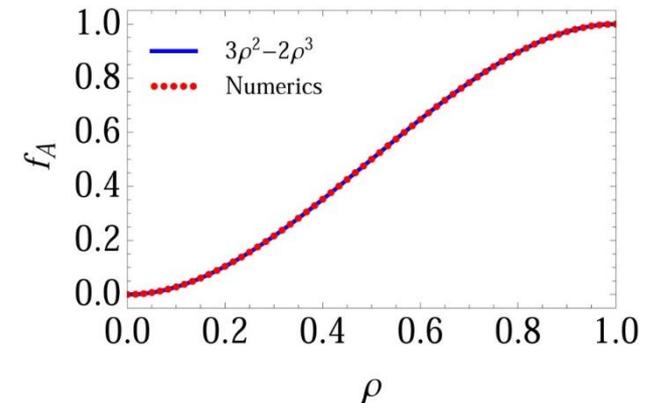
## ❖ Quasiparticle description:

Low-energy spectrum of excitations of some systems are governed by effectively conformal theories, when both temperatures are low. *Bernard, Doyon '16*

...so the highest lying parts of the spectrum are not populated.

➤ Universal formula should be valid in ballistic regimes of actual electronic systems. Correlation functions too? Lattice model expectations?

extremized for  $z_j = t\sqrt{\ell - t}$



$$\begin{aligned} d_R(z_j, x) = & \log \left[ (1 + \pi^2 T_R^2 \tilde{z}_j^2) \cosh(2\pi T_R(x - \ell)) - (1 - (\pi T_R \tilde{z}_j)^2) \cosh(2\pi T_R(t - x)) \right] + \\ & \log \left[ (1 + \pi^2 T_L T_R \tilde{z}_j^2) \cosh(\pi(t T_L - t T_R + 2T_R x)) \right. \\ & \left. + (\pi^2 T_L T_R \tilde{z}_j^2 - 1) \cosh(\pi(t(T_L + T_R) - 2T_R x)) \right] - \frac{1}{2} \log(16\pi^8 T_L^2 T_R^6 \tilde{z}_j^4) \end{aligned}$$

# MATCHING GEODESICS - RESULTS

❖ Extremization:

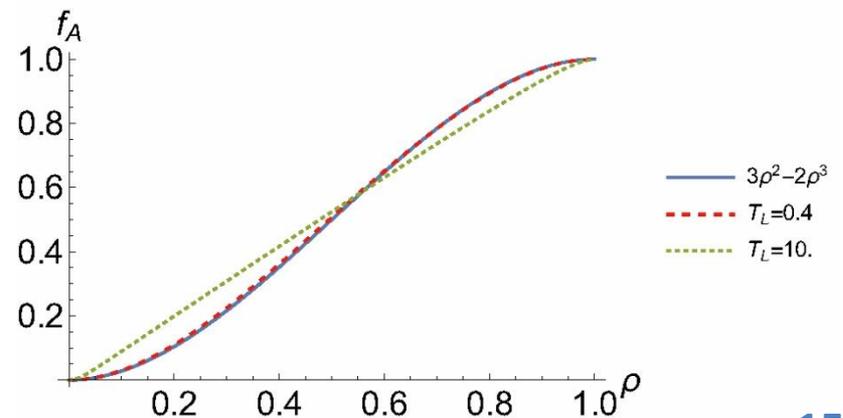
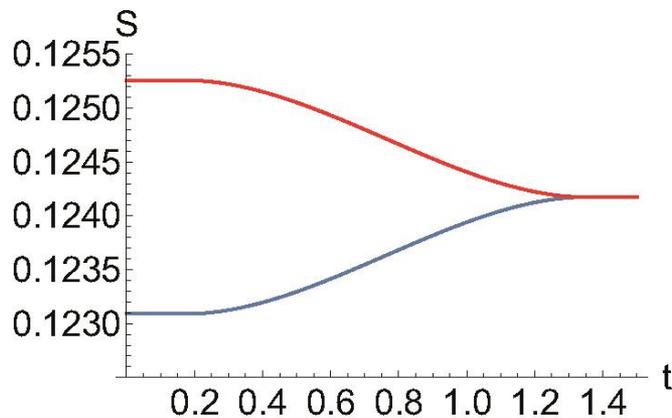
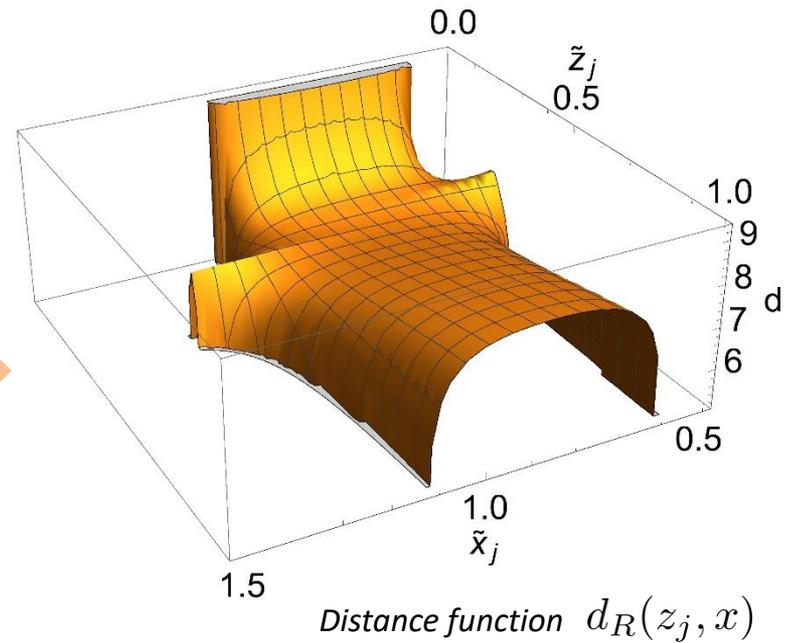
$$\partial_{\tilde{z}_j} d_R = 0, \quad \partial_x d_R = 0$$

→ Numerical methods to solve non-linear algebraic equations.

❖ The distance function turns imaginary outside of some region.

(if one boundary point becomes null or timelike-separated from the joining point)

❖ Argument from Kruskal diagram → Exclude solutions with  $\tilde{z}_j > \tilde{z}_H$



# VELOCITY IN ENTANGLEMENT GROWTH

*Liu, Suh '13*

*Li, Wu, Wang, Yang '13*

*Hartman, Maldacena '13*

- After a global quench, the entanglement entropy exhibits quadratic growth:

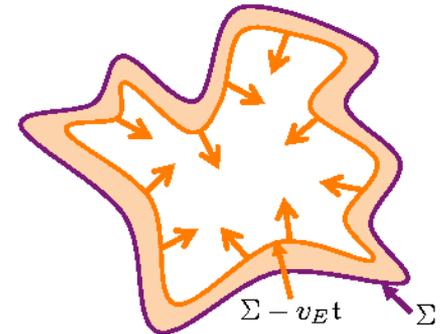
$$\Delta S(t) \propto t^2 + \dots$$

- Followed by a universal linear growth regime where

$$\Delta S(t) = v_E s_{eq} A_\Sigma t + \dots$$

- The velocity  $v_E$  depends on the final equilibrium state. In the case of an AdS-RN black hole,

$$v_E = \frac{\sqrt{d}(d-2)^{\frac{1}{2}-\frac{1}{d}}}{(2(d-1))^{1-\frac{1}{d}}} \longleftarrow \text{Tsunami Velocity}$$



- **Butterfly velocity:** Speed of propagation of chaotic behavior in the boundary theory:

$$v_B = \sqrt{\frac{d}{2(d-1)}}$$

*Shenker, Stanford '13*

*Roberts, Stanford, Susskind '14*

$$W_x(t) = e^{-iHt} W_x e^{iHt}$$

For an operator local on the thermal scale, defined on a Tensor Network

Bound between these velocities:

$$1 \geq v_B \geq v_E$$

# BOUNDS IN VELOCITIES

Rangamani, Rozali, Vincart-Emard '17

- Average Velocity**

Average entropy increase rate:

$$v_{av} \equiv \frac{\Delta S}{\Delta t} = \frac{L}{4G\ell} \log \left( \frac{T_L \sinh(\pi\ell T_R)}{T_R \sinh(\pi\ell T_L)} \right)$$

This quantity is bounded, although it can be arbitrarily large:

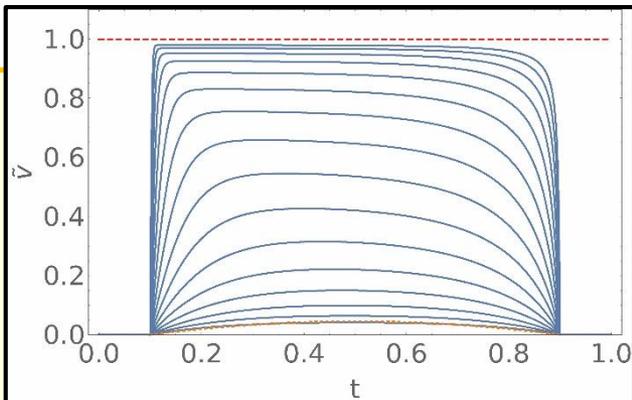
$$\lim_{\ell \rightarrow \infty} v_{av} = \frac{L}{4G} \pi |T_R - T_L|$$

Normalized by the entropy density of the final state, we find

$$|\tilde{v}_{av}| \leq \left| \frac{T_R - T_L}{T_R + T_L} \right| \leq 1 \quad \leftarrow \text{To compare with } \Delta S(t) = v_E s_{eq} A_\Sigma t + \dots,$$

where  $\tilde{v}_{av} \equiv v_{av}/s_{eq}$

➤ When normalized in a **physical** way, we get a similar bound as 2d entanglement tsunamis or local quenches.

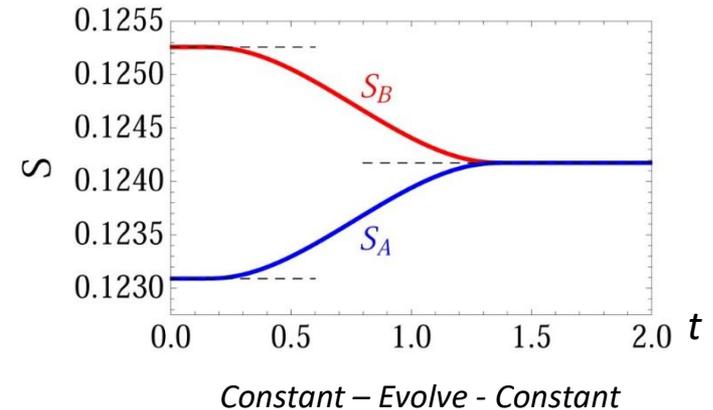


- Momentary Velocity**

$$\tilde{v} \equiv \frac{1}{s_{eq}} \frac{dS(\ell, t)}{dt} \leq 1$$

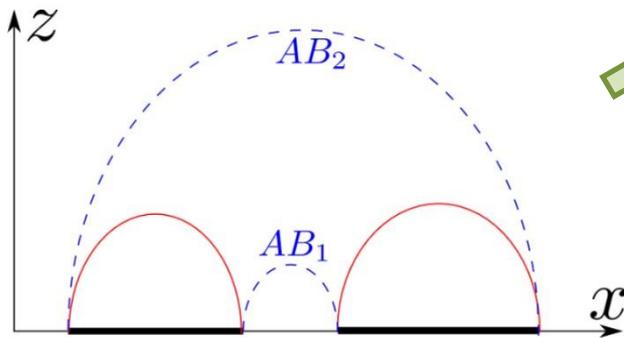
Numerically, we still find this bound.

➤ Interpretation: The shockwave seems to take the role that the entanglement tsunami had for a global quench.



# $n > 2$ DISCONNECTED INTERVALS

*Hubeny, Rangamani, Takayanagi '07*



Two physical configurations for calculating the entanglement entropy.

- Choose the minimal possible configuration:

$$S(AB) = \min \{ S(A) + S(B), S_{AB_1} + S_{AB_2} \}$$



Phase transitions! **Configurations = Phases**

Entanglement entropies are required to satisfy certain inequalities

✓ Subadditivity:  $S(AB) \leq S(A) + S(B)$

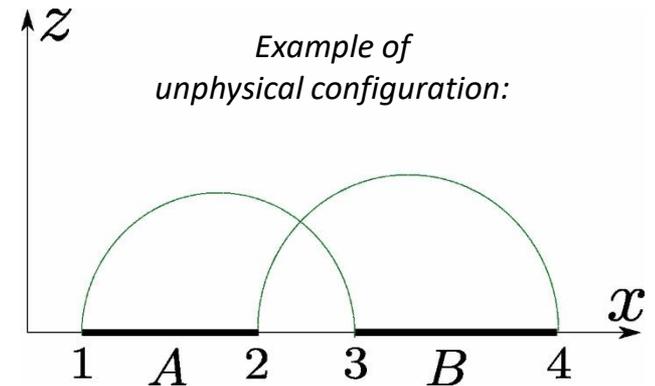
*Araki, Lieb '70*

✓ Triangle:  $S(AB) \geq |S(A) - S(B)|$

- Similar concepts with  $n > 2$  intervals?

*Mirabi, Tanhayi, Vazirian '16*

*Bao, Chatwin-Davies '16*



Example of unphysical configuration:

- When enumerating the possible phases, we must exclude those with curves intersecting (*unphysical* phases)

# PHYSICAL INTERVAL PHASES

$$S(AB) = S(A) + S(B) \Leftrightarrow \begin{pmatrix} 1 \rightarrow 2 \\ 3 \rightarrow 4 \end{pmatrix} \text{ "disconnected phase" } \img alt="Two separate semi-circles above intervals [1,2] and [3,4] on a line." data-bbox="718 168 798 192"/>$$

$$S(AB) = S(AB_1) + S(AB_2) \Leftrightarrow \begin{pmatrix} 1 \rightarrow 4 \\ 2 \rightarrow 3 \end{pmatrix} \text{ "connected phase" } \img alt="A single large semi-circle above [1,4] and a smaller one above [2,3] on a line." data-bbox="718 238 798 282"/>$$

## Unphysical configurations

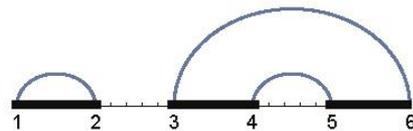
- Do not yield lowest values for the entanglement entropy.
- In a time-dependent case, the co-dimension one surface spanned would become null or timelike.

*Headrick, Takayanagi '07*  
*Hubeny, Maxfield, Rangamani, Tonni '13*

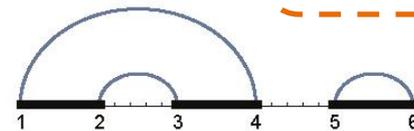
Phase 1:



Phase 2:



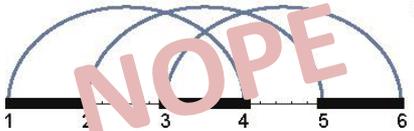
Phase 3:



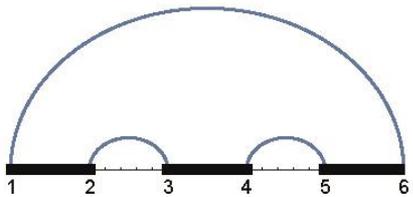
$$\binom{N}{2n} = \frac{N!}{(2n)!(N-2n)!}$$

ways to join intervals

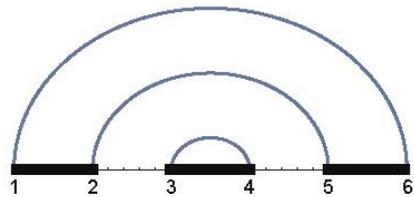
Phase 4:



Phase 5:



Phase 6:



# GENERALIZED INEQUALITIES

## n=3 case

- Strong Subadditivity inequality:

*Lieb, Ruskai '73*

✓ Time dep. case

$$S(AB) + S(BC) - S(ABC) - S(B) \geq 0$$

*Headrick, Takayanagi '07*

- A different inequality, which was proven for the holographic prescription:

$$S(AB) + S(BC) - S(A) - S(C) \geq 0$$

- Monogamy of mutual information == Negativity of tripartite information:

$$I_3(A : B : C) \equiv S(A) + S(B) + S(C) - S(AB) - S(BC) - S(AC) + S(ABC) \leq 0$$

*A. Wall '12*

✓ Time dep. case

## n>3 cases

- For n=5 intervals (A, B, C, D, E), this generalizes to 5 inequalities.

- Negativity of n-partite information:

$$I_n(A_1 : A_2 : A_3 : \dots : A_n) \equiv \sum_{i=1}^n S(A_i) - \sum_{i<j}^n S(A_i \cup A_j) + \sum_{i<j<k}^n S(A_i \cup A_j \cup A_k)$$

$$\mp \dots + (-1)^n S(A_1 \cup A_2 \cup \dots \cup A_n),$$

- Proposed inequalities:

*Alishahiha, Mozaffar, Tanhayi '14*  
*Mirabi, Tanhayi, Vazirian '16*

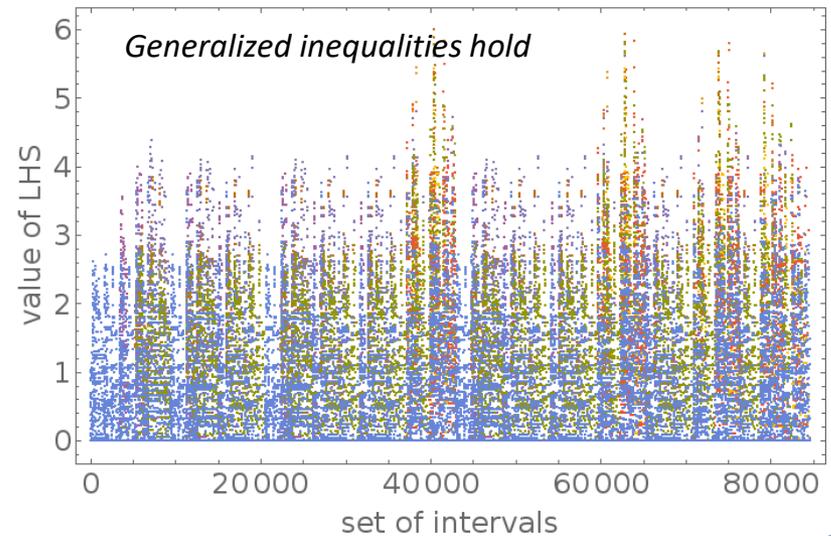
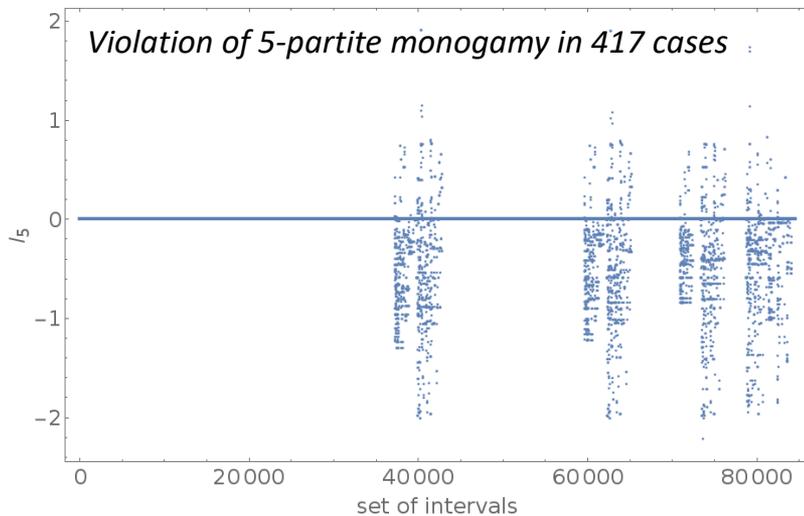
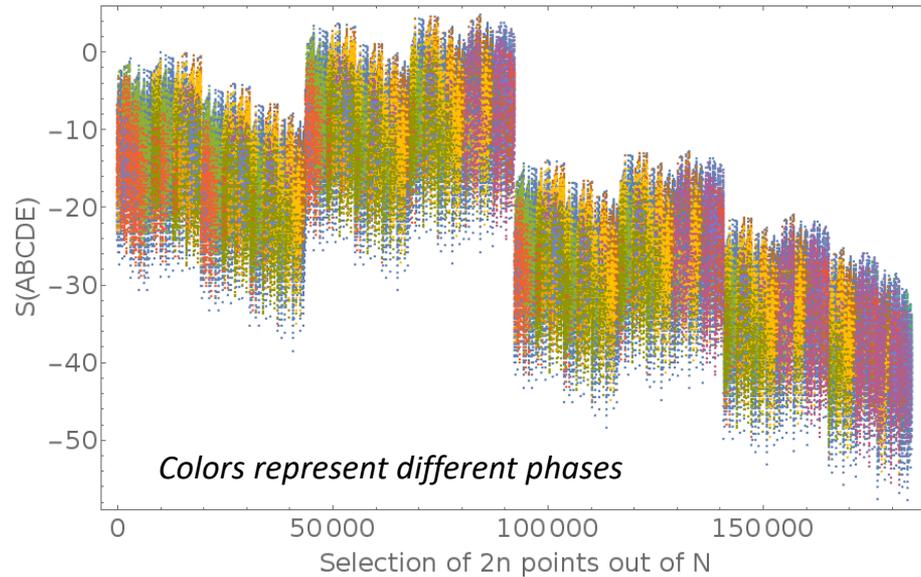
$$\left[ \begin{array}{l} I_4(A : B : C : D) \geq 0 \\ I_5(A : B : C : D : E) \leq 0 \end{array} \right.$$

...which do **not** hold in holographic setups.

*Hayden, Headrick, Maloney '11*

# RESULTS FOR $n=5$ INTERVALS

- 42 physical phases
- 20 boundary points
- 184756 possible unions
- 84579 not totally disconnected



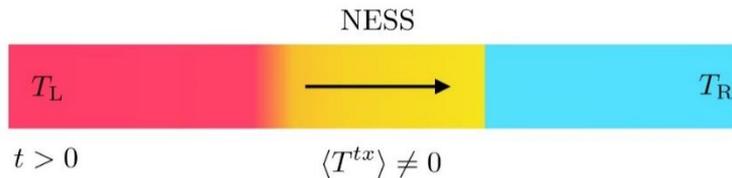
# HIGHER DIMENSIONS

- ❖ A solution was found in the hydrodynamic regime
- ❖ A similar solution in the holographic setup confirmed it
- ❖ An inconsistency between results and thermodynamics was found

*Bhaseen, Doyon, Lucas '15*

*Amado, Yarom '15*

*Spillane, Herzog '15*



- The higher-dimensional case is more physically relevant and interesting.

**Assuming that the dual-shock solution is valid approximately:**

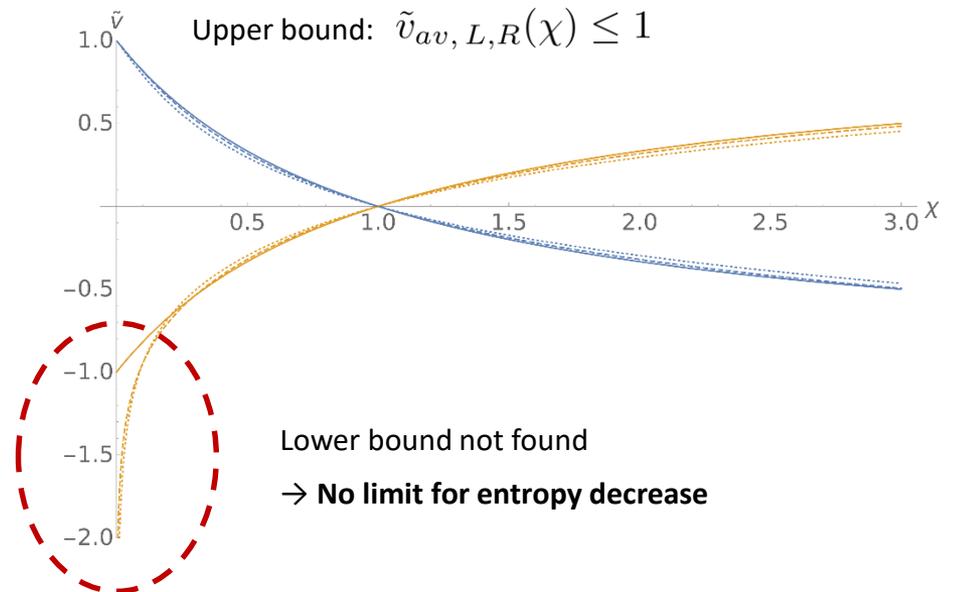
The shockwaves move with different velocities:

$$u_L = \frac{1}{d-1} \sqrt{\frac{\chi + d - 1}{\chi + \frac{1}{d-1}}}, \quad u_R = \sqrt{\frac{\chi + \frac{1}{d-1}}{\chi + d - 1}}.$$

Statements about velocity bounds, similar to

$$0 \leq |v_{av}| \leq \frac{L}{4G} \pi |T_R - T_L|$$

can be derived for higher dimensions.



# CONCLUSIONS AND REMARKS

- ❖ **Universal** steady state, described by boosted black brane.
- ❖ Entanglement Entropy measures information flow.
- ❖ Mutual Information grows **monotonically** in time.
- ❖ Entanglement Entropy decrease and increase rates are **bounded**.
- ❖ Shockwaves mimic the entanglement **tsunami**.
- ❖ **Inequalities** are satisfied and violated, confirming expectations.

Universal formula:

$$f_A(\rho) \simeq 3\rho^2 - 2\rho^3$$

## Outlook 1:

- This bulk metric is vacuum – Null Energy Condition is satisfied.

Will time-dependent bulk spacetimes that violate NEC still satisfy the inequalities?

*Callan, He, Headrick '12  
Caceres, Kundu, Pedraza, Tangarife '13*

## Outlook 2:

- The low temperature regime of a lattice model can be approximated by a CFT thermal state

Can our simple universal evolution be observed in Tensor Network calculations?

*Bohrdt, Mendl, Endres, Knap '16*

*Thank you for your attention!*